

Geometry of the Full 2-Body Problem

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Example of a FBP: asteroid pairs.



Context

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□ Others will speak about dynamical systems aspects of these problems, including transport rates, etc.

Important Tools

- □ For mechanical systems with symmetry, some of the tools are:
 - Momentum maps, ie conserved quantities
 - *Reduction*, shape space
 - Stability and the energy-momentum method
 - Geometric phases

Restricted Problems

 \Box **Restricted** means that one part of the system evolves in the field of another part of the system; Examples are

- Spherical pendulum on a Merry-Go-Round (the pendulum dynamics does not affect the rotation of the Merry-Go-Round)
- Fluid flow on a rotating earth (the fluid does not affect the Earth's rotation)
- Restricted 3-body problem

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- Restricted 3-body problem
- □ We typically handle restricted problems by the theory of *moving systems*.

Restricted 3-body Problem

- consider the *planar case*—the *spatial case* is similar
- *Kinetic energy* (wrt inertial frame) in rotating coordinates:

$$K(x, y, \dot{x}, \dot{y}) = \frac{1}{2} \left[(\dot{x} - \omega y)^2 + (\dot{y} + \omega x)^2 \right]$$

 \circ **Lagrangian** is K.E. – P.E., given by

$$L(x, y, \dot{x}, \dot{y}) = K(x, y, \dot{x}, \dot{y}) - V(x, y); \quad V(x, y) = -\frac{1 - \mu}{r_1} - \frac{\mu}{r_2}$$

• Euler-Lagrange equations:

$$\ddot{x} - 2\omega \dot{y} = -\frac{\partial V_{\omega}}{\partial x}, \qquad \ddot{y} + 2\omega \dot{x} = -\frac{\partial V_{\omega}}{\partial y}$$

where the *effective potential* is

$$V_{\omega} = V - \frac{\omega^2 (x^2 + y^2)}{2}$$

Effective potential

• Equations for the third body are those of a *particle moving in an effective potential plus a magnetic field* (Jacobi, Hill, etc)



Effective Potential



Level set shows the Hill region

More Tools of the Trade

□ Geometric mechanics provides a general theory for (usually) nonrestricted problems: mechanical systems with symmetry. Eg, notions of amended potiential, relative equilibria, stability by the energy-momentum method, variational integration algorithms (symplectic integrators), etc.



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- \Box The full 2-body problem has a SE(3) symmetry and corresponding conserved quantities
- $\Box \ Configuration \ Manifold: \ Q = SE(3) \times SE(3)$
- \Box The **shape space** Q/G gives the **system shape** and plays an important role in reduction theory.
- \Box Lots of work by many people, as in Dan Scheeres talk

- $\square Material points in a reference configuration X_i; i = 1 for body 1 and i = 2 for body 2$
- \Box Points in the current configuration x_i .

 \Box Given a configuration

 $((A_1, r_1), (A_2, r_2)) \in SE(3) \times SE(3),$

the *material* and *spatial* points are related by

 $x_1 = A_1 X_1 + r_1$ and $x_2 = A_2 X_2 + r_2$



Lagrangian equals kinetic minus potential energy:

$$\begin{split} L(A_1, r_1, A_2, r_2, \dot{A}_1, \dot{r}_1, \dot{A}_2, \dot{r}_2) \\ &= \frac{1}{2} \int_{\mathcal{B}_1} \|\dot{x}_1\|^2 d\mu_1(X_1) + \frac{1}{2} \int_{\mathcal{B}_2} \|\dot{x}_2\|^2 d\mu_2(X_2) + \int_{\mathcal{B}_1} \int_{\mathcal{B}_2} \frac{G d\mu_1(X_1) d\mu_2(X_2)}{\|x_1 - x_2\|} \\ &= \frac{m_1}{2} \|\dot{r}_1\|^2 + \frac{1}{2} \langle \Omega_1, I_1 \Omega_1 \rangle + \frac{m_2}{2} \|\dot{r}_2\|^2 + \frac{1}{2} \langle \Omega_2, I_2 \Omega_2 \rangle \\ &+ \int_{\mathcal{B}_1} \int_{\mathcal{B}_2} \frac{G d\mu_1(X_1) d\mu_2(X_2)}{\|A_1 X_1 - A_2 X_2 + r_1 - r_2\|}. \end{split}$$

□ Here, for instance, $\Omega_1 = A_1^{-1}\dot{A}_1$ is the body angular velocity of the first body (with the usual identification of 3 × 3 skew matrices with vectors.

Goal: Reduce by overall translations and rotations and bring the machinery of geometric mechanics to bear.

 \square SE(3) acts by the diagonal left action on Q: $(A, r) \cdot (A_1, r_1, A_2, r_2) = (AA_1, Ar_1 + r, AA_2, Ar_2 + r).$

 \Box Momentum map

$$\mathbf{J}: TQ \to \mathfrak{se}(3)^*$$

is the total linear and angular momentum. \Box Shape space Q/G: one copy of SE(3); coordinatized by the relative attitude $A = A_1^{-1}A_2$ and relative position $R = A_2^T(r_1 - r_2)$.

Some Reduction Theory

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 \Box In this case, one gets a natural identification

 $\alpha_A: (TQ)/G \to T(Q/G) \times \tilde{\mathfrak{g}}$

where $\tilde{\mathfrak{g}} = (Q \times \mathfrak{g})/G$ is the **associated bundle**. Similarly for the Hamiltonian side of the story.

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- \Box First of all, **shape space** is given by

$$X = (G \times G)/G \cong G,$$

where the map $\pi: Q = G \times G \to G$ is given by

$$x = \pi(g_1, g_2) = g_1^{-1}g_2.$$

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 \Box This has picked out one of the bodies as special.

 \Box This gives rise to the identification

$\left(T(G\times G)\right)/G\cong G\times\mathfrak{g}\times\mathfrak{g}$

where we map the class of $(g_1, \dot{g}_1, g_2, \dot{g}_2)$ to (x, w, ξ_2) , where $x = g_1^{-1}g_2$ and $\xi_2 = g_2^{-1}\dot{g}_2$ as above and where $w = \dot{x}x^{-1}$.

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 \Box Thus, by general theory, the equations of motion will reduce to equations for the variables (x, w, ξ_2) .

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- □ One gets, therefore, two sets of equations, one set of second order equations for the shape space variables $x \in Q/G$ and another set of equations for the variables in \tilde{g} . These equations are called the *Lagrange-Poincaré* equations.
- □ The equations correspond to breaking up the variational principle into two parts: one for horizontal variations (*Lagrange* part of the equations) and one for vertical variations (*Poincaré* part of the equations).

□ One can work this all out quite explicitly for the general case of $Q = G \times G$, where, say, G = SE(3) and for Lagrangians of the form

$$L(g_1\xi_1, g_2\xi_2) = \frac{1}{2} \left[\operatorname{Tr}(K_1\xi_1^T\xi_1) + \operatorname{Tr}(K_2\xi_2^T\xi_2) \right] - V(g_1^{-1}g_2).$$

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- □ Similarly, one gets the Hamiltonian version of the equations, the reduced Poission structure, etc.

Systematic Structures

- □ For numerics as well as analysis of stability of relative equilibria (analog of the libration points), the variational and Hamiltonian structures are useful.
- Previous works guessed these structures and missed the variational structure altogether. Using reduction, one derives them in a simple and natural way, one gets the Jacobi integrals naturally, etc.
- Extra symmetries give extra conserved quantities and further reductions.

- □ Restricted (as in restricted 3-body problem) simple case already exhibits the basic ejection and collision dynamics
- \Box Point mass moving in the *xy*-plane under the gravitational field of a uniformly rotating elliptical body, without affecting its uniform rotation.

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- \Box Point mass moving in the *xy*-plane under the gravitational field of a uniformly rotating elliptical body, without affecting its uniform rotation.
- **Equations of motion** relative to a rotating Cartesian coordinate frame and appropriately normalized:

$$\ddot{x} - 2\dot{y} = \frac{\partial V}{\partial x}$$
 and $\ddot{y} + 2\dot{x} = \frac{\partial V}{\partial y}$,

where

$$V(x,y) = \frac{1}{\sqrt{x^2 + y^2}} + \frac{1}{2}(x^2 + y^2) + U_{22};$$

and where

$$U_{22} = \frac{3C_{22} \left(x^2 - y^2\right)}{\left(x^2 + y^2\right)^{5/2}}$$

 \Box The coefficient C_{22} is the *ellipticity*.

 \Box Jacobi integral: $J = \frac{1}{2} \left(\dot{x}^2 + \dot{y}^2 + \dot{z}^2 \right) - V.$

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□ *Moving systems approach* gives, as in the RCTBP, the Lagrangian and Hamiltonian structure and Jacobi integral.

□ Lagrangian (kinetic minus potential energy) written in the rotating system and with angular velocity normalized to unity, is

$$L = \frac{1}{2} [(\dot{x} - y)^2 + (x + \dot{y})^2 + \dot{z}^2] - U(x, y, z).$$

where

$$U(x, y, z) = -\frac{1}{r} - U_{22}.$$

Euler-Lagrange equations produce the previous equations and the Legendre transformation gives the Hamiltonian structure, the Jacobi integral, etc.

- □ The Jacobi integral (energy) is an indicator of the type of global dynamics possible.
- □ For energies above a threshold, $E > E_S$, corresponding to symmetric saddle points, movement between the *realm* near the asteroid (*interior realm*) and away from the asteroid (*exterior realm*) is possible. For energies $E \leq E_S$, no such movement is possible.

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- \Box As in the CRTBP, motion between realms is mediated by phase space *tubes*.
- □ General theory allows us to transition what we learned in the CRTBP to this case.

Phase space in each realm organized further into different *resonance regions*, connected via *lobes*.



□ Poincaré sections in the different realms, U_1 and U_2 are linked by tubes in the phase space. Under the Poincaré map f_1 on U_1 , a trajectory reaches an **exit**, the last Poincaré cut of a tube before it enters another realm. The map f_{12} takes points in the exit of U_1 to the **entrance** of U_2 . The trajectory then evolves under the action of the Poincaré map f_2 on U_2 .

□ See Shane's talk and the material on the FBP website for further information.

Selected References

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