Design of Low Energy Space Missions using Dynamical Systems Theory

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► Main Theme

• how to use dynamical systems theory of 3-body problem in low energy trajectory design.

Background and **Motivation**:

- NASA's Genesis Discovery Mission.
- A Low Energy Tour of Jupiter's Moons.
- ► Restricted 3-Body Problem.
- ► Main Results.

► Ongoing Work.

- Low Thrust Trajectories in a Multi-Body Environment.
- Parking a Satellite near an Asteroid Pair.

Motivation: Genesis Discovery Mission

► Genesis spacecraft

• collected solar wind sample from a L_1 halo orbit,

• **returned** them to Earth.

► Halo orbit, transfer/ return trajectories in rotating frame.



Motivation: Genesis Discovery Mission

Designed using dynamical systems theory (Barden, Howell, and Lo).

- ► Followed **natural dynamics**, little propulsion after launch.
- **Return-to-Earth portion** utilized heteroclinic dynamics.



Motivation: Petit Grand Tour of Jupiter's Moons

- Construct a low energy trajectory to visit several moons in one mission.
- ▶ Instead of flybys, can **orbit each moon for any duration**.
- ▶ NASA is considering a Jupiter Icy Moon Orbiter (**JIMO**).



Design Strategy: Patched 3-Body Solutions

- Jupiter-Ganymede-Europa-SC 4-body system approximated as 2 coupled 3-body systems
- ► **3-body solutions** of each 3-body systems are linked in right order to construct orbit with desired itinerary.
- ▶ Try to **minimize** ΔV at each transfer patch point.
- ▶ Initial solution refined in **4-body model**.
- ▶ 3-body solutions offer a large class of low energy trajectories.



Planar Circular Restricted 3-Body Problem

\triangleright 2 main bodies

- Total mass normalized to 1: $m_J = \mu$, $m_S = 1 \mu$.
- Rotate about center of mass, angular velocity normalized to 1.
- ► Choose **rotating** coordinate system with origin at center of mass, 2 main bodies fixed at $(-\mu, 0)$ and $(1 \mu, 0)$.



Equilibrium Points

▶ Equations of motion for SC are

$$\ddot{x} - 2\dot{y} = -\frac{\partial U}{\partial x}, \qquad \ddot{y} + 2\dot{x} = -\frac{\partial U}{\partial y},$$

where $U(x, y) = -\frac{x^2 + y^2}{2} - \frac{1 - \mu}{r_s} - \frac{\mu}{r_j}.$

► Five equilibrium points:

• 3 **unstable** collinear equilbrium points, L_1, L_2, L_3 .

• 2 equilateral equilibrium points, L_4, L_5 .



Hill's Realm

► Energy integral: $E(x, y, \dot{x}, \dot{y}) = (\dot{x}^2 + \dot{y}^2)/2 + U(x, y).$

- ► E can be used to determine (**Hill's**) **realm** in position space where SC is energetically permitted to move.
- ► Effective potential: $U(x,y) = -\frac{x^2+y^2}{2} \frac{1-\mu}{r_s} \frac{\mu}{r_j}$.



Hill's Realm

► To fix energy value E is to fix height of plot of U(x, y). Contour plots give 5 cases of Hill's realm.



The Flow near L_1 and L_2

- For energy value just above that of L_2 , Hill's realm contains a "neck" about $L_1 \& L_2$.
- SC can make **transition** through these equilibrium realms.
- ► 4 types of orbits:

periodic, asymptotic, transit & nontransit.



Invariant Manifold as Separatrix

- Asymptotic orbits form 2D invariant manifold tubes in 3D energy surface.
- ▶ They separate transit and non-transit orbits:
 - Transit orbits are those inside the tubes.
 - Non-transit orbits are those outside the tubes.



Invariant Manifold as Separatrix

▶ Invariant Manifold Tubes associated with periodic orbits about L_1 , L_2 control ballistic capture and escape.



Heteroclinic Connection

- ► Found **heteroclinic connection** between pair of periodic orbits.
- ▶ Found a large class of **orbits** near this (homo/heteroclinic) *chain*.
- \triangleright SC can follow these *channels* in rapid transition.



Existence of Transitional Orbits

- Main Theorem: For any admissible itinerary, e.g., (..., X, J; S, J, X, ...), there exists an orbit whose whereabouts matches this itinerary.
- ► Can even specify **number of revolutions** the comet makes around Sun & Jupiter (plus $L_1 \& L_2$).

▶ **3-Body** trajectories much richer than **2-body** trajectories.



Numerical Construction of Orbits

- Developed procedure to construct orbit with prescribed itinerary.
- \blacktriangleright Example: An orbit with itinerary $(\mathbf{X}, \mathbf{J}; \mathbf{S}, \mathbf{J}, \mathbf{X})$.



Construction of $(\mathbf{M}, \mathbf{X}; \mathbf{M}, \mathbf{I}, \mathbf{M})$ Orbits

- ▶ Invariant mfd. **tubes** $(S \times I)$ separate transit/nontransit orbits.
- ▶ Red curve (S^1) (Poincaré cut of L_2 unstable manifold). Green curve (S^1) (cut of L_1 stable manifold).
- ▶ Any point inside the intersection region Δ_M is a (X; M, I) orbit.



Construction of (M,X;M,I,M) Orbits

- ▶ The desired orbit can be constructed by
 - Choosing appropriate **Poincaré sections** and
 - linking invariant **manifold tubes** in right order.



Petit Grand Tour of Jupiter's Moon

- ▶ Petit Grand Tour can be constructed similarly
 - Approximate 4-body system as 2 nested **3-body systems**.
 - Choose appropriate **Poinaré section**.
 - Link invariant **manifold tubes** in right order.
 - Refine initial solution in **4-body model**.



Petit Grand Tour of Jupiter's Moons (Planar Model)

Used invariant manifolds

to construct trajectories with interesting characteristics:

- Petit Grand Tour of Jupiter's moons.
 1 orbit around Ganymede. 4 orbits around Europa.
- A ΔV nudges the SC from
 Jupiter-Ganymede system to Jupiter-Europa system.

▶ Instead of **flybys**, can orbit several moons for **any duration**.







Extend from Planar Model to Spatial Model



Look for Natural Pathways to Bridge the Gap

- **Tubes** of two 3-body systems **may not intersect** for awhile. May need large ΔV to "jump" from one tube to another.
- ► Look for **natural pathways** to bridge the gap
 - between z_0 where tube of one system (Ganymede) enters and z_2 where tube of another system exits (into Europa realm) by "hopping" through phase space (z_1) .



Transport in Phase Space via Tube & Lobe Dynamics

► By using

- **tubes** of rapid transition that connect realms
- **lobe dynamics** to hop through phase space,

New tour only needs $\Delta V = 20$ m/s (50 times less).



Lobe Dynamics: Mixed Phase Space
 Poincaré section reveals mixed phase space:
 resonance regions and

• "chaotic sea".



Transport between Regions via Lobe Dynamics

- **Invariant manifolds** divide phase space into resonance regions.
- ► Transport between regions can be studied via **lobe dynamics**.





Transport between Regions via Lobe Dynamics

- Segments of **unstable** and **stable** manifolds form **partial barriers** between regions R_1 and R_2 .
- \blacktriangleright $L_{1,2}(1), L_{2,1}(1)$ are **lobes**; they form a **turnstile**.
 - In one iteration, only points from R₁ to R₂ are in L_{1,2}(1)
 only points from R₂ to R₁ are in L_{2,1}(1).
- ▶ By studying pre-images of $L_{1,2}(1)$, one can find efficient way from R_1 to R_2 .



Hopping through Resonaces in Low Energy Tour

- Guided by lobe dynamics, hopping through resonances (essential for low energy tour) can be performed.
- ► To get SC captured by secondary (m_2) , need to decrease **semi-major axis** passing through **resonances**.



Tube/Lobe Dynamics: Transport in Solar System

- ► To use **tube** dynamics/**lobe** dynamics of **spatial** 3-body problem to **systematically** design low-fuel trajectory.
- Part of our program to study transport in solar system using tube and lobe dynamics.



Low Thrust Trajectories in a Multi-Body Environment

► Incorporation of **low thrust**.

▶ Design to take best advantage of **natural dynamics**.

► See Shane D. Ross.



Spiral out from Europa



Europa to lo transfer

Parking a Satellite near an Asteroid Pair

- ► Find stable periodic/quasi-periodic orbits for SC to observe binary as it orbits the Sun.
 - Model for asteroid pair:
 sphere and rigid body (3 connected masses).
 - Model for SC motion: binary in **relative equilibrium**.
- ► See Gabern, Koon, and Marsden [2004].



